

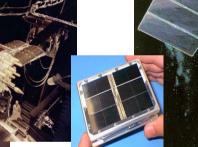
LADE E COST



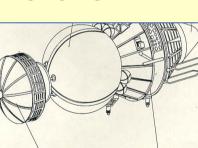


Space Junk

Jonathan McDowell







Power Supply





















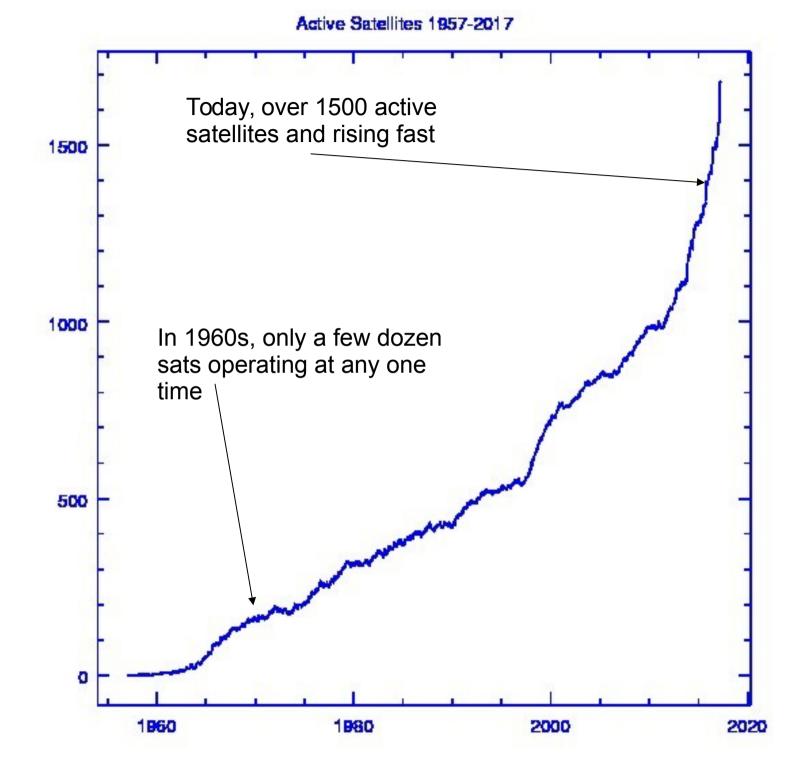




About 1500 satellites currently operating

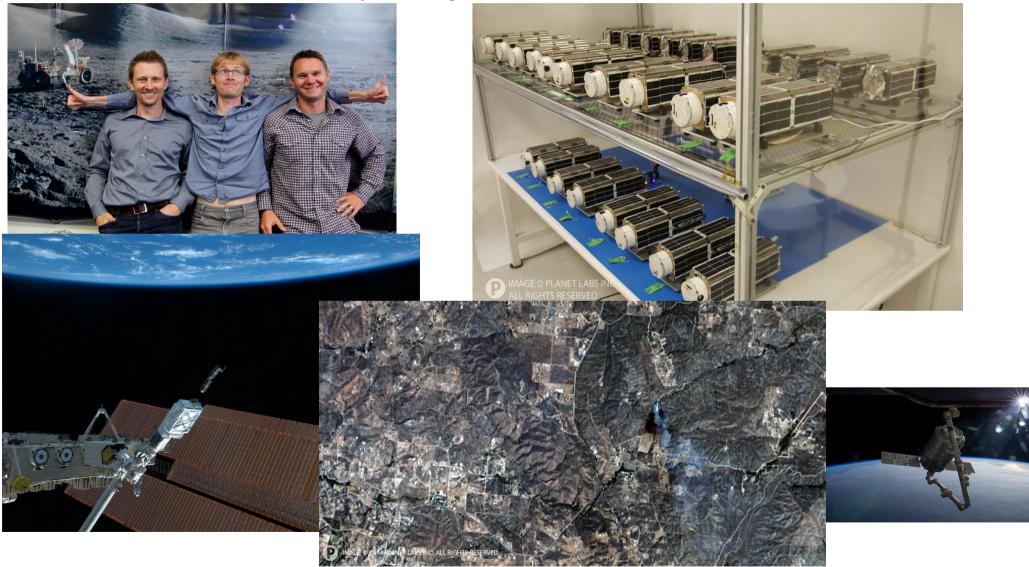
Some in low orbit skimming just outside the atmosphere, mostly going from pole to pole

Some In 'geostationary orbit' in a ring high above the equator



2013: CUBESAT EXPLOSION!

99 Cubesats launched Jun 2003-Feb 2013 by 63 organizations in 20 countries 120 Cubesats launched Mar 2013 – Feb 2014 by 57 organizations in 18 countries (Cumulative: 219 Cubesats by 108 orgs in 28 countries)

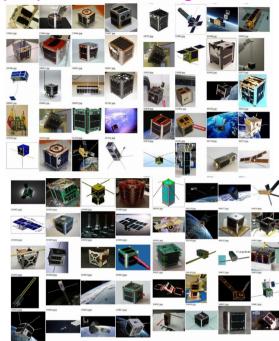


Chris, Will and Robbie left NASA to found PlanetLabs in a San Fran office building – 71 satellites launched since 2013, first big Cubesat constellation

The Cubesat Explosion: STATISTICS 2017 April

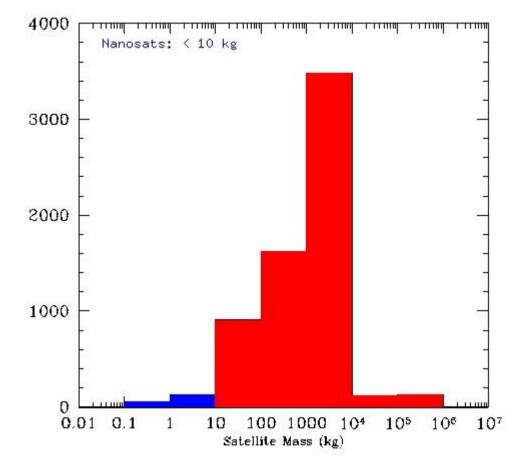
Cubesat Launches 150 2 Aboard ISS Working? 291 Dead 57 Reentered 179 Sep Fail 5 Returned 8 Launch Fail 69 611 Total 100 50 Ô 2005 2010 2015

http://planet4589.org

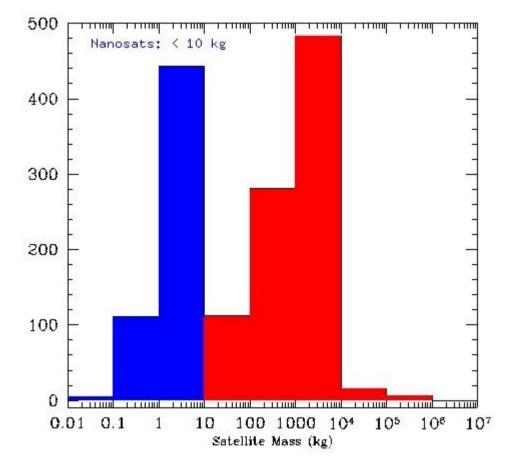


TOTAL 38 COUNTRIES:

USA 478 Japan 27 Germany 10 China 10 Denmark 9 Singapore 6 Netherlands, Italy 5 S Korea, Belgium, Spain, India 4 Canada, UK, Brazil, Israel, Norway, Peru 3 Russia, Switzerland, Lithuania, Turkey, Vietnam, Ecuador , Argentina 2 Ukraine, France, Kazakhstan, Emirates, Uruguay, S Africa, Algeria, Poland, Pakistan, Colombia, Romania, Hungary, Estonia 1



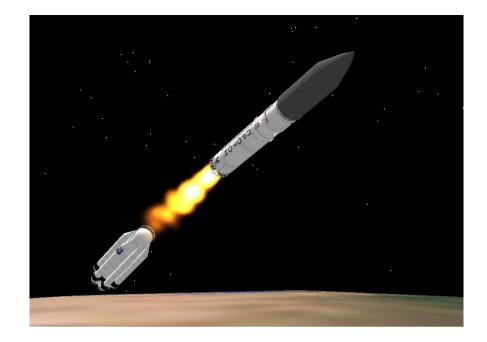
Payload Mass Distribution 1957-2009

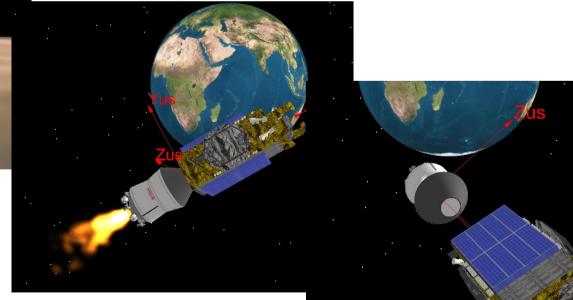


Payload Mass Distribution 2010-2017

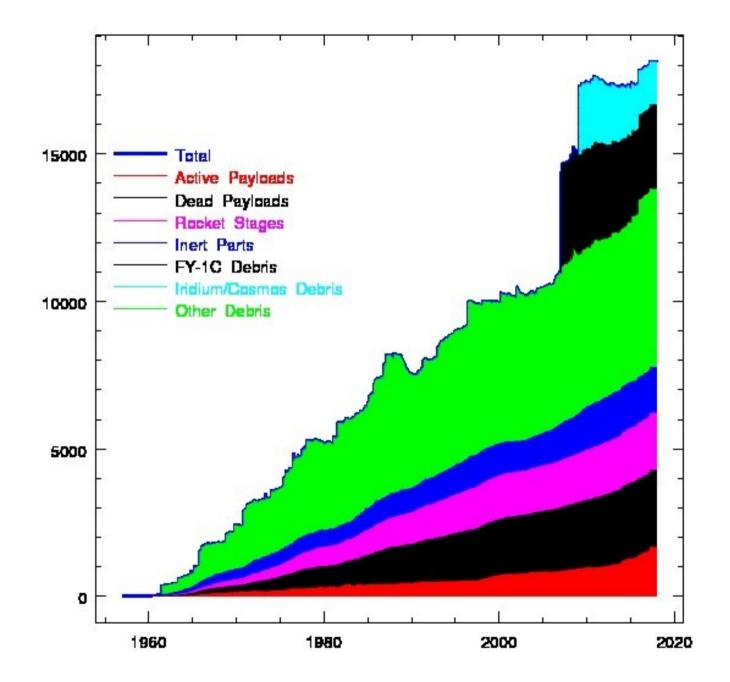


A typical satellite launch ends up with at least two objects in orbit – the satellite and the last piece ("stage") of the rocket that got it there

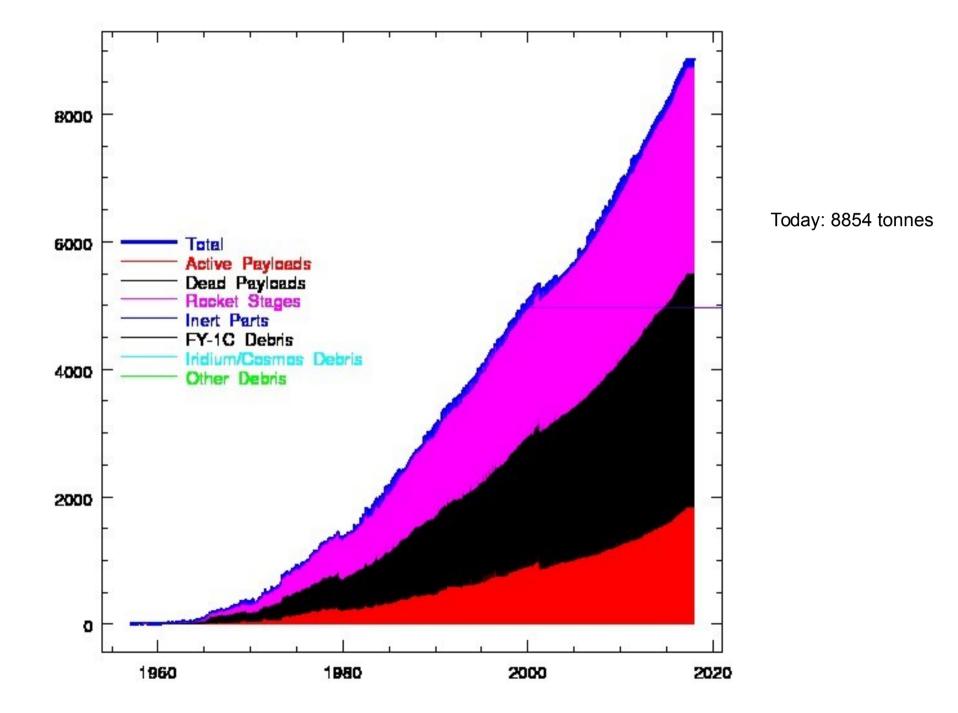




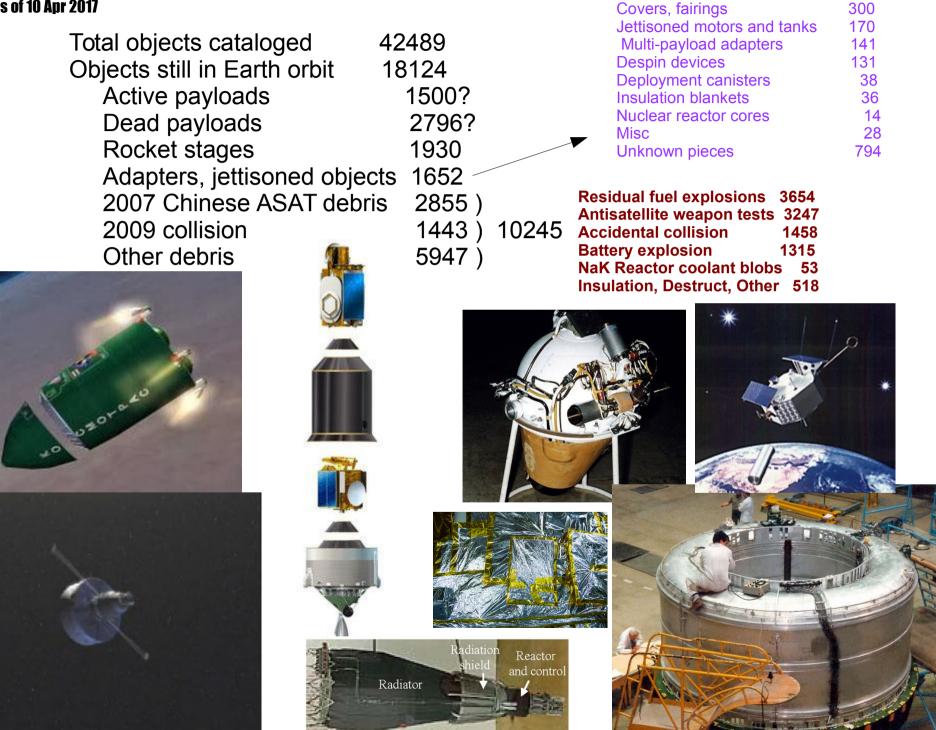
The Growth of Space Junk



Space Junk - mass in metric tons

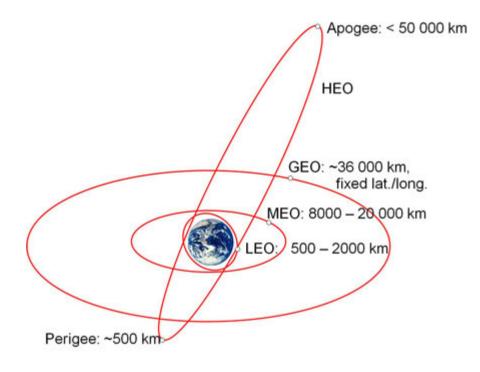


A Census of Space Debris as of 10 Apr 2017



300

Orbitography





Low Earth Orbit



Earth surface has r = 6378 km Space Station has height 400 km, so r = 6778 km

This corresponds to v = 7.67 km/s or v = 17158 mph - quite fast!! At 400 km, orbital period is 92.5 minutes Kepler in sensible units

Newton's law of gravitation

$$V = mc^2 \left(\frac{r}{R_G}\right)^-$$

Keplerian velocity

$$v/c = \sqrt{R_G/r}$$

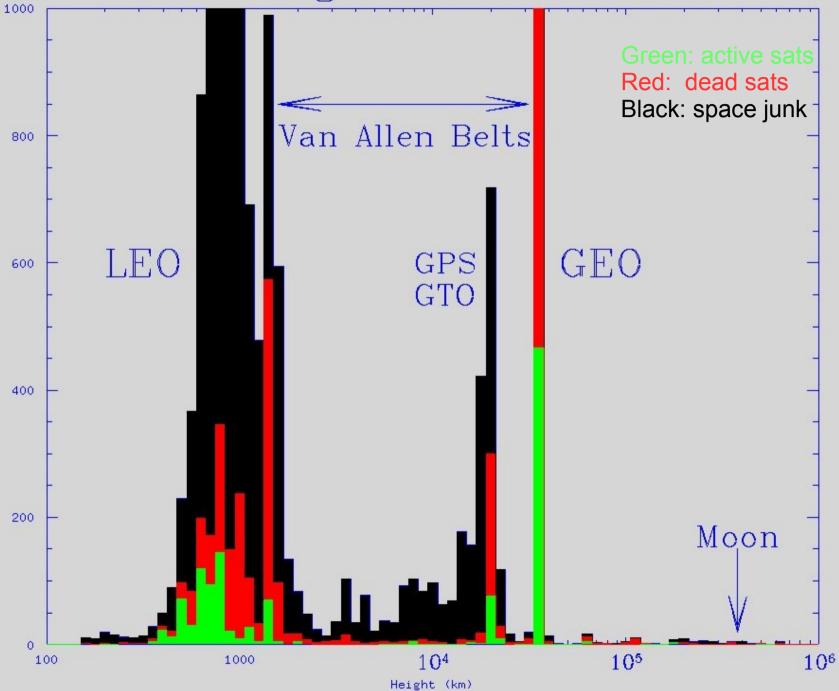
Kepler's third law

$$T = \frac{2\pi r}{c} \sqrt{r/R_G}$$

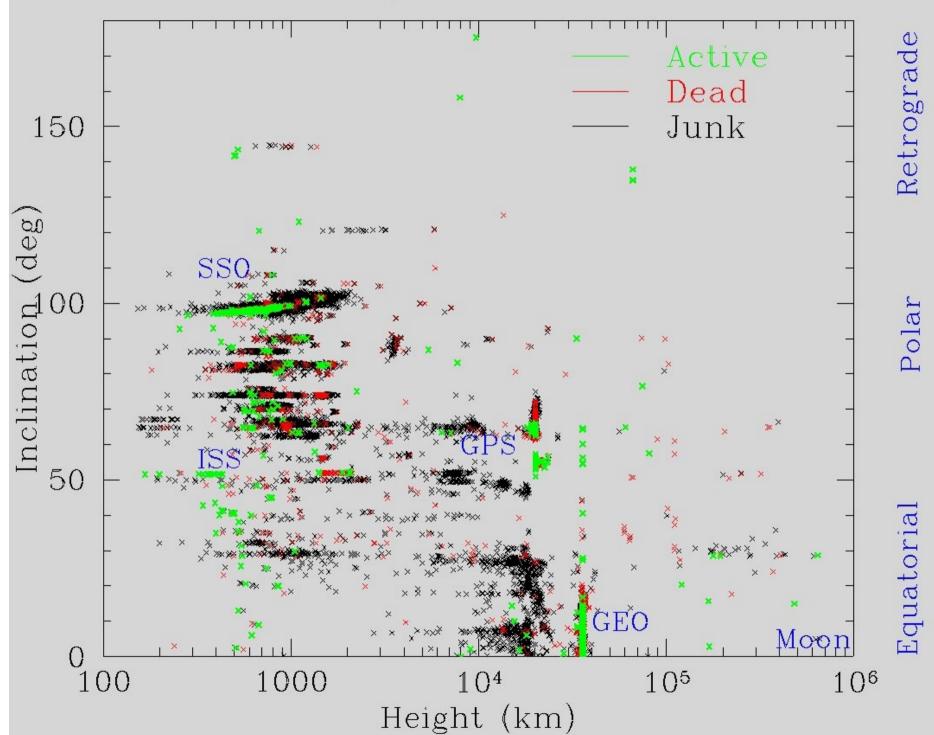
$$R_G = \frac{GM}{c^2}$$

is the gravitational radius of a central mass M.

How high are satellites?

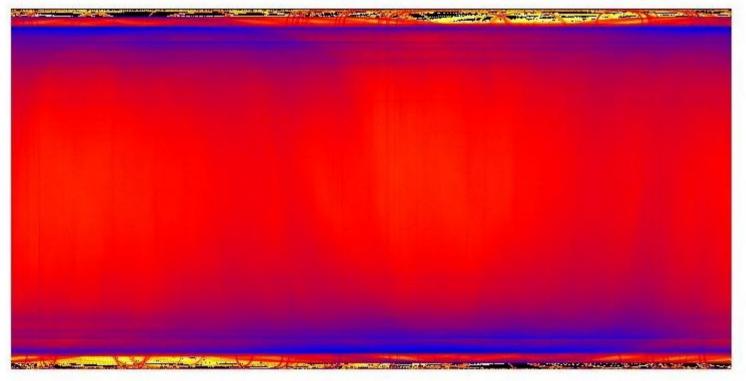


A Map Of Earth Orbit



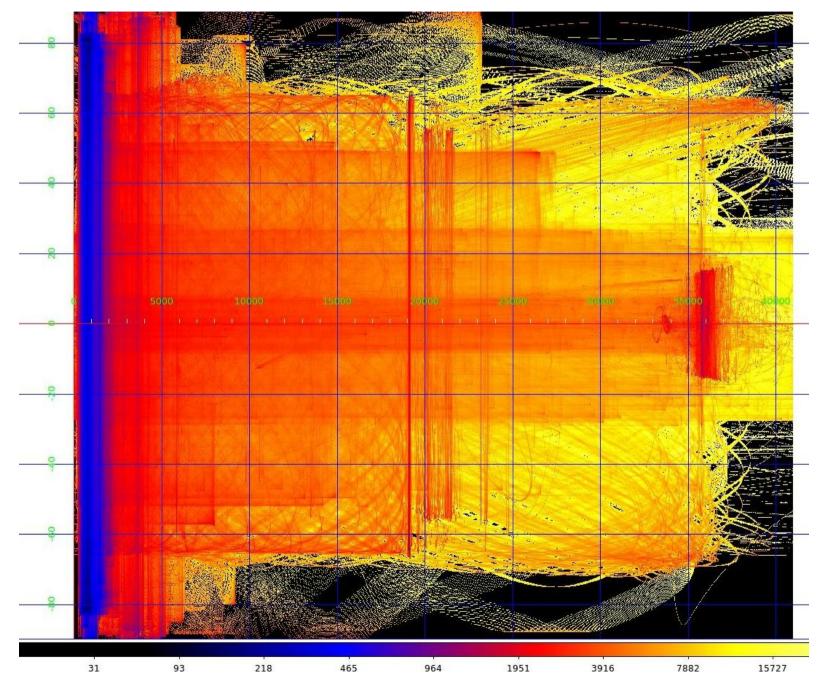
Mean free path in 800 km LEO: RA vs latitude

Calculated positions for 18000 satellites each minute for 45 day period to obtain average number density in 1 x 1 deg bins extending +-50 km in radial direction



	Contract of the second				Prov		1		
1	4	10	21	44	89	178	359	717	

Here is the longitude-averaged MFP as a function of height (x axis) versus latitude (y axis) Note the GEO and GPS belts and the horizontal structures



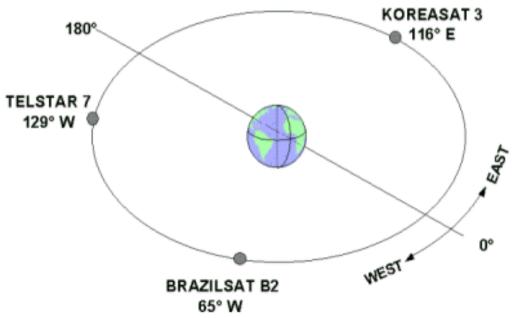
GEO: Geostationary Earth Orbit

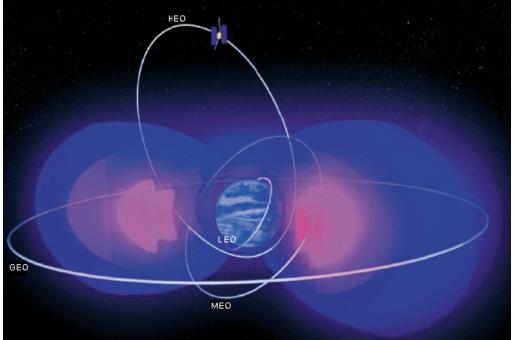
Consider a satellite whose orbit goes around the Earth's equator Just outside the atmosphere it takes 1 ½ hours to go round the planet Far out, at the distance of the Moon it takes a month to go round Inbetween there is some height at which it takes exactly 23 hr 56 min

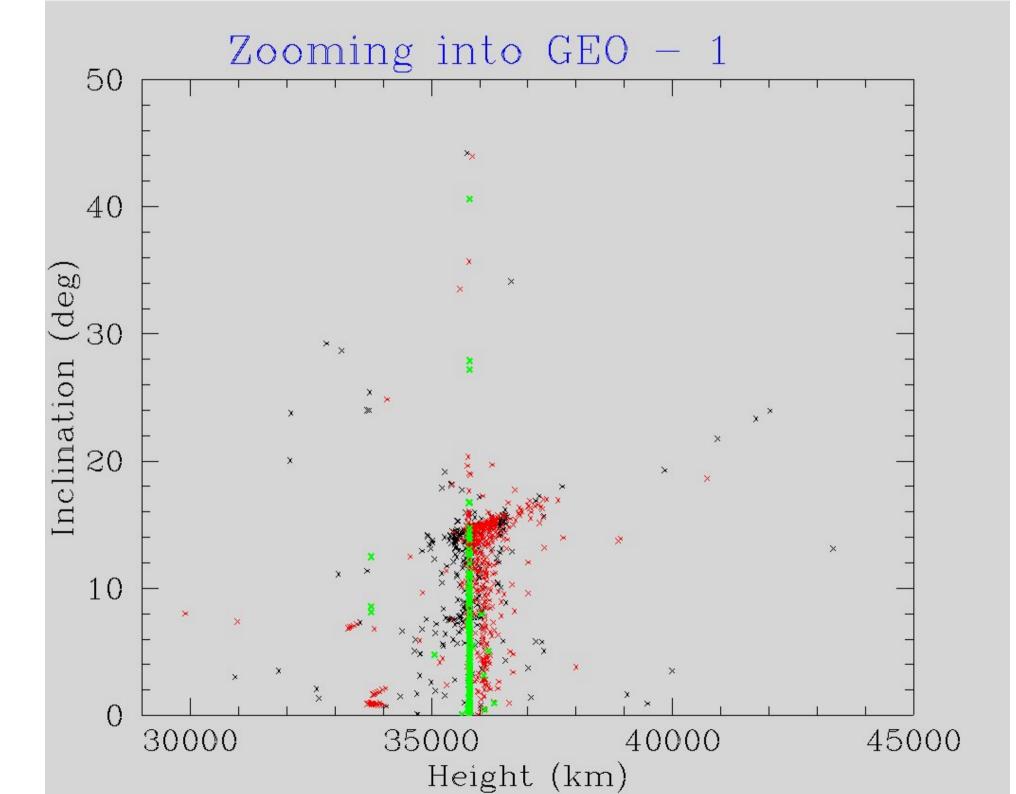
Meanwhile, the Earth spins underneath it, also taking 23 hr 56 min to complete one full rotation

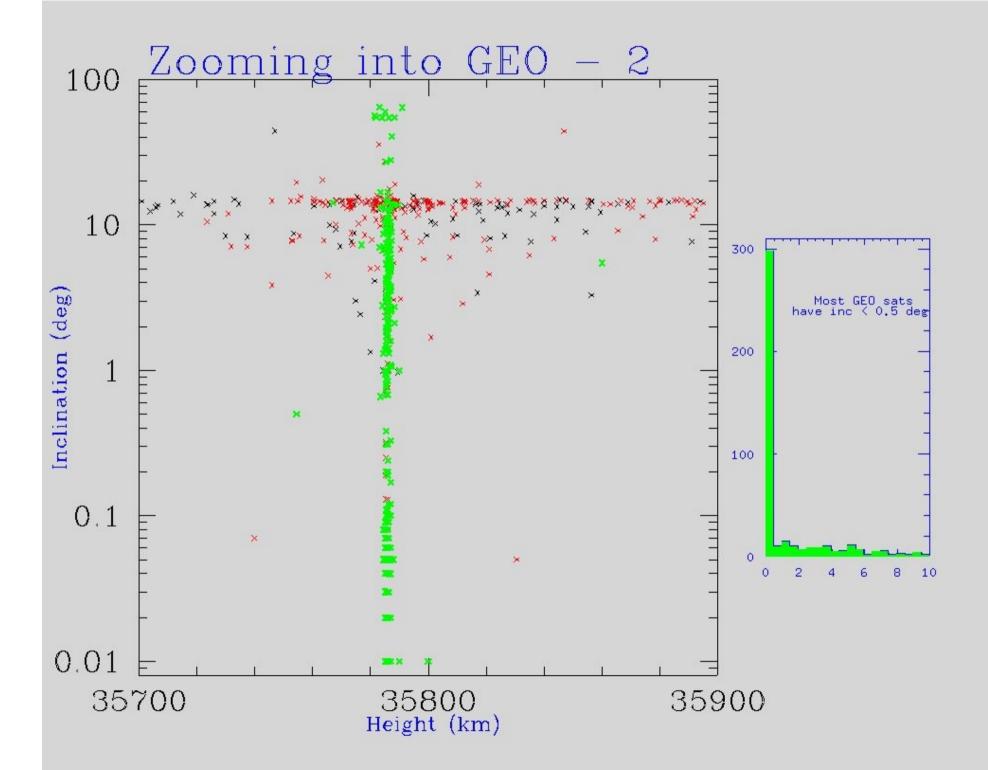
So the satellite stays above the same point on the equator!

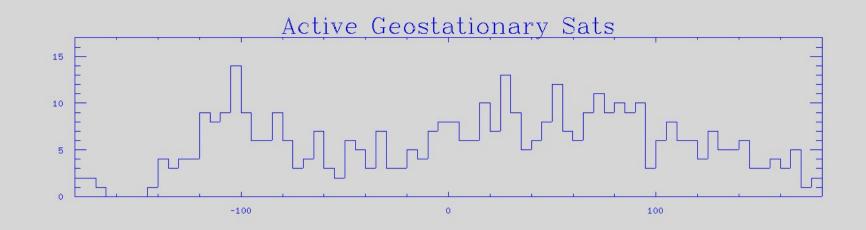
Kepler's Third Law lets us calculate the magic height: 35787 km above the Earth's surface (about 23000 miles)

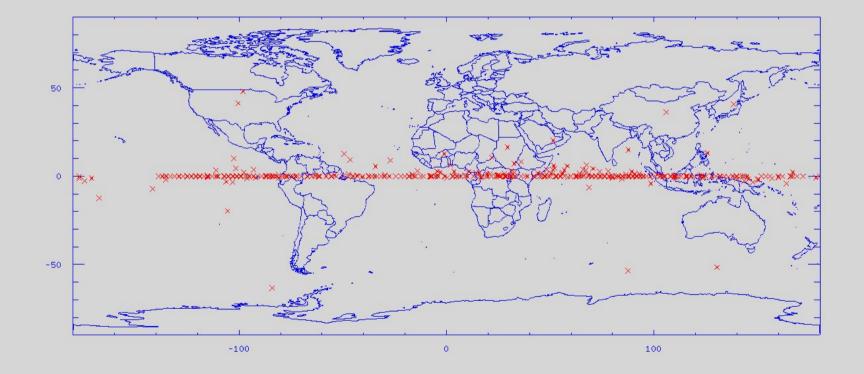












Special orbits due to the oblate Earth

Actually we left something out of our math: the Earth is NOT ROUND! It's a little squashed at the poles (polar radius is 22 km smaller than at equator) Every time a sat goes over the poles, it gets less of a tug; over the equator it gets more. This twists the orbit – makes it rotate in space. By picking the orbit cleverly you can make the twist do something useful! We consider the first term (J2) in the spherical harmonic expansion of the potential This gives first order corrections to the orbital elements (node, arg of peri.)

- varying linearly in time

The first order orbital rotation rate for a flattened object of equatorial radius R_s

$$f_p = \frac{(3\pi J_2) \left(R_s/a\right)^{\frac{7}{2}} (1-e^2)^{-2}}{T_s}$$

where T_s is the surface orbital period

$$T_s = \frac{2\pi R_s}{c} \sqrt{\frac{R_s}{R_G}}$$

The node and perigee then rotate at

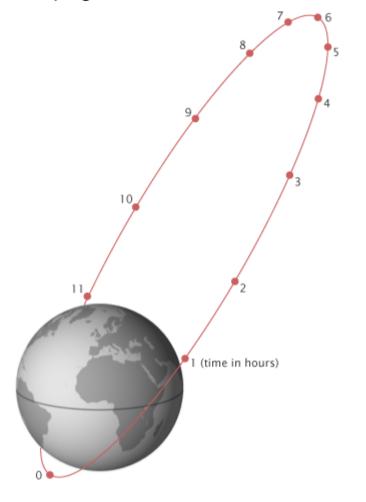
$$\dot{\omega} = f_P \left(2 - \frac{5}{2} \sin^2 i \right)$$

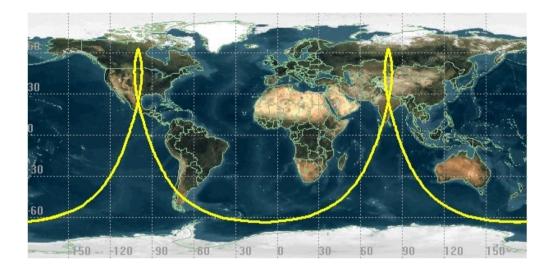
and

 $\dot{\Omega} = f_P \cos i$

First let's consider the perigee rotation.

Suppose you have an elliptical inclined orbit with apogee over the northern hemisphere





The oblateness will tend to rotate the orbit in its orbital plane – soon the apogee will be in the south.

But by careful choice of inclination we can set the term $2 - 5/2 \sin^2 i$ to zero: $\sin^2 i = 4/5$ or i = 63.43 degrees

and lock in the latitude of apogee

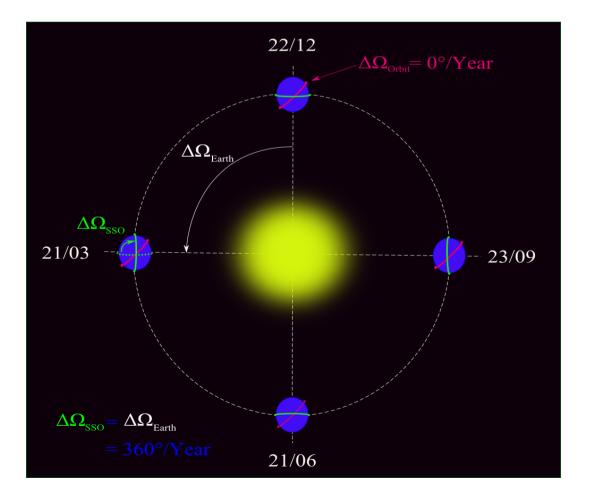
the "Molniya orbit" (usually T = 12 hr)

SSO: Sun Synchronous Orbit

Now let's make the ascending node term do something useful.

In perfect Keplerland, the orbit plane (and so, RA of ascending node) is fixed in inertial space. Hence, the angle between the orbit plane and the Earth-Sun line changes as the Earth orbits the sun, by about 1 degree a day.

The correct choice of altitude and inclination can induce an opposite motion in the orbit plane to keep the orbit-normal/Earth-Sun-line angle fixed.

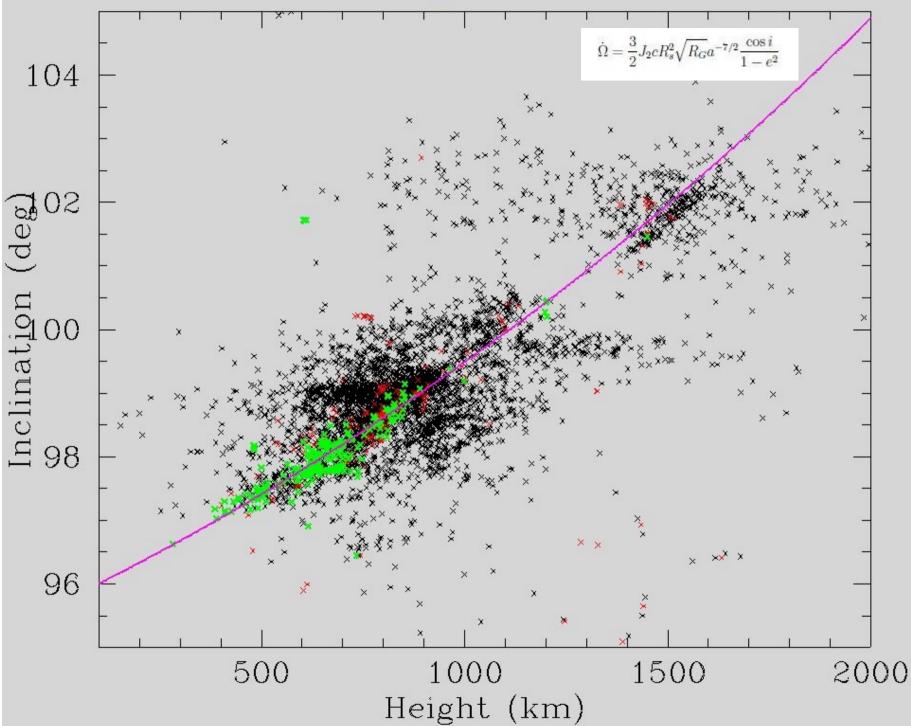


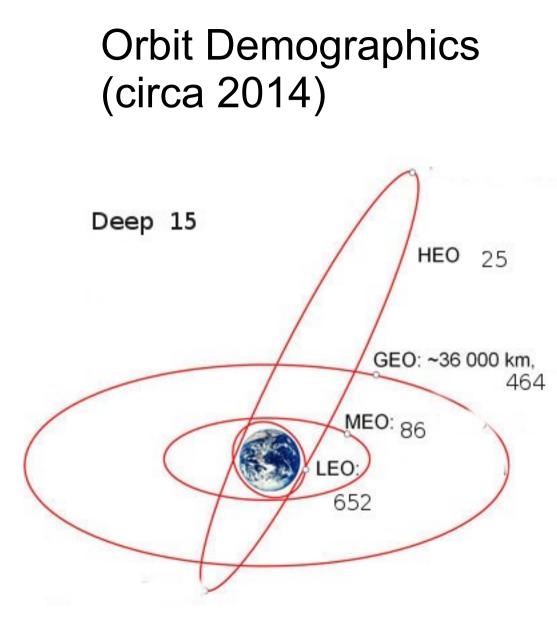
The magenta colored orbit is what you get for a perfect sphere Earth

It stays fixed in space so in August (in this particular case) it is facing the sun – the satellite orbits over the dawn/dusk line - but in May the orbit is edge on to the sun, orbiting noon to midnight.

The green colored orbit is SSO, turning so it's always facing the Sun

Zooming into SSO





Overall statistics:

A	ctive	Dead	Junk
LEO	652	1512	10327
MEO	86	262	758
HEO/GTO	25	151	1562
GEO	464	518	291
Deep	15	51	62

*Most satellites are either in LEO or GEO

Special cases: LEO - SSO 282 247 - others 370 1265

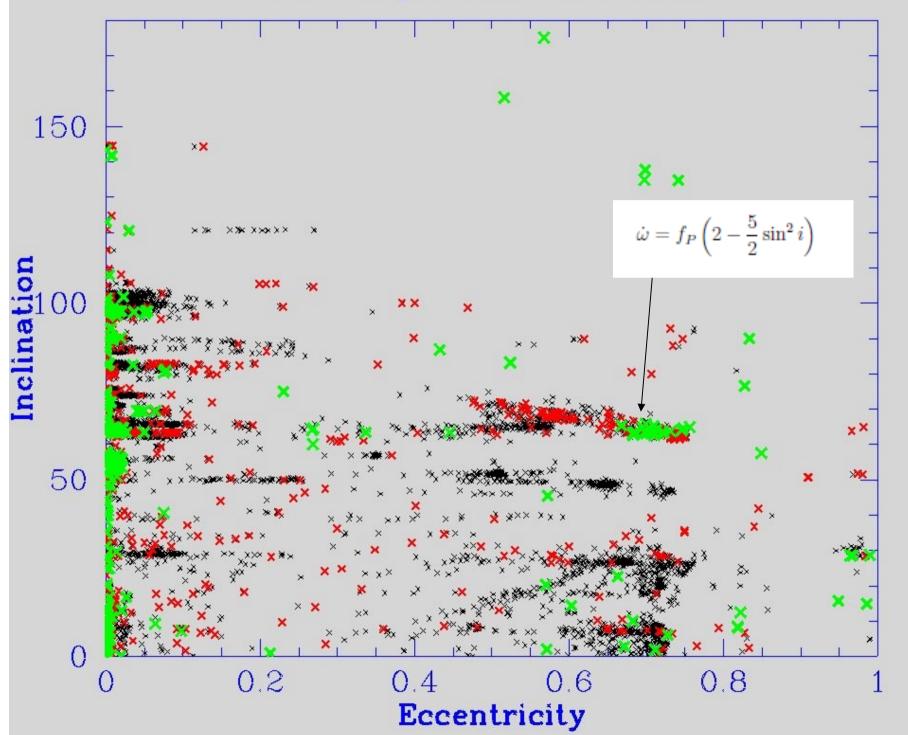
2 247 5173 0 1265 4625

SSO is a very specific orbit, has almost half the LEO sats – and most of the debris

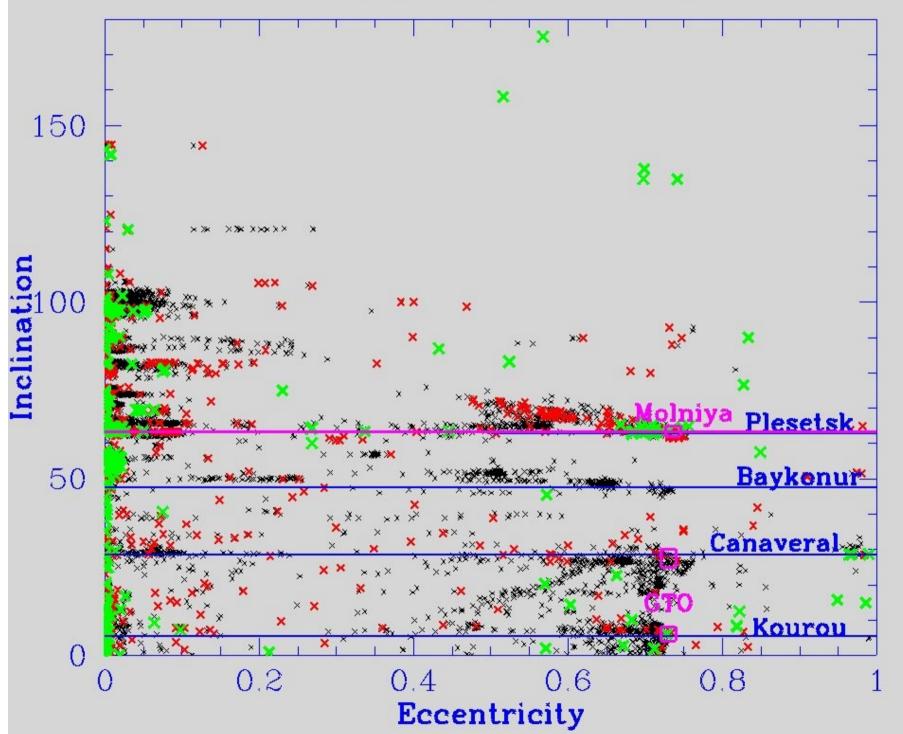
GEO			
Stationary	452	83	14
Graveyard	5	187	92
Drift	7	233	167
Other	1	22	57

Only 36 percent of dead GEO sats are in the graveyard

How Elliptical vs. How Polar



How Elliptical vs. How Polar



Collisions

In LEO: Most likely collision over the poles, where SSO orbits in different planes intersect.

Example: 2009 Iridium/Strela collision

SSN 24946 Iridium 33, launched 1997 for Iridium LLC. Dry mass 556 kg

SSN 22675 Strela-2M 56, launched 1993 for Russian MoD. Dry mass 800 kg Codename Kosmos-2251; retired 1995

Collision Feb 10 1656 UTC at

97.9E 72.5N Alt. 776 km over the Siberian Arctic

Sat 1: 7.465 km/s 12 deg E of N Sat 2: 7.470 km/s 14 deg S of E

Relative velocity 11.64 km/s

KE of Strela in Irid frame 54 GJ (Comparison: 1 ton truck @ 100mph = 1 MJ) Some damage was done...

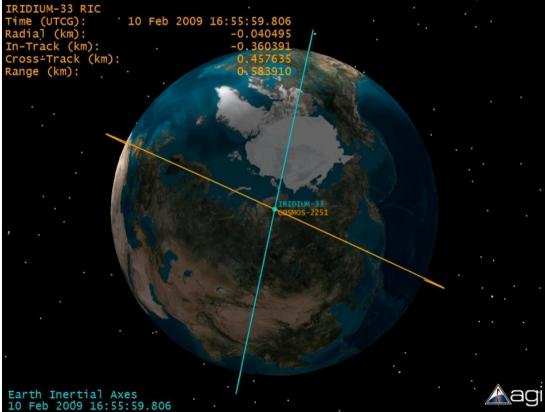
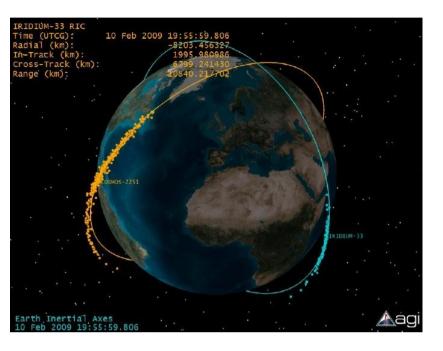
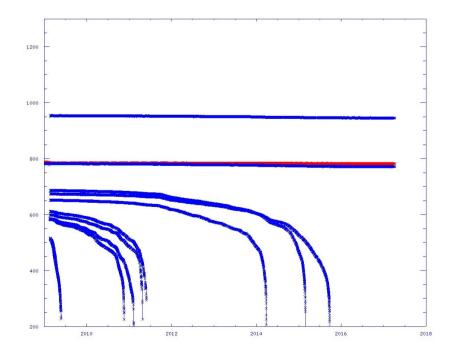


Figure from Kelso 2009



Height versus time for selected Strela-2M debris



3 hours post collision (image from Kelso 2009) the debris spreads out along the orbit of each satellite

(compare meteor streams along comet orbits)

Eventually debris objects spread in RA due to differential orbital precession to make a shell

Also spread in altitude due to varying A/m ratio and hence drag coefficient

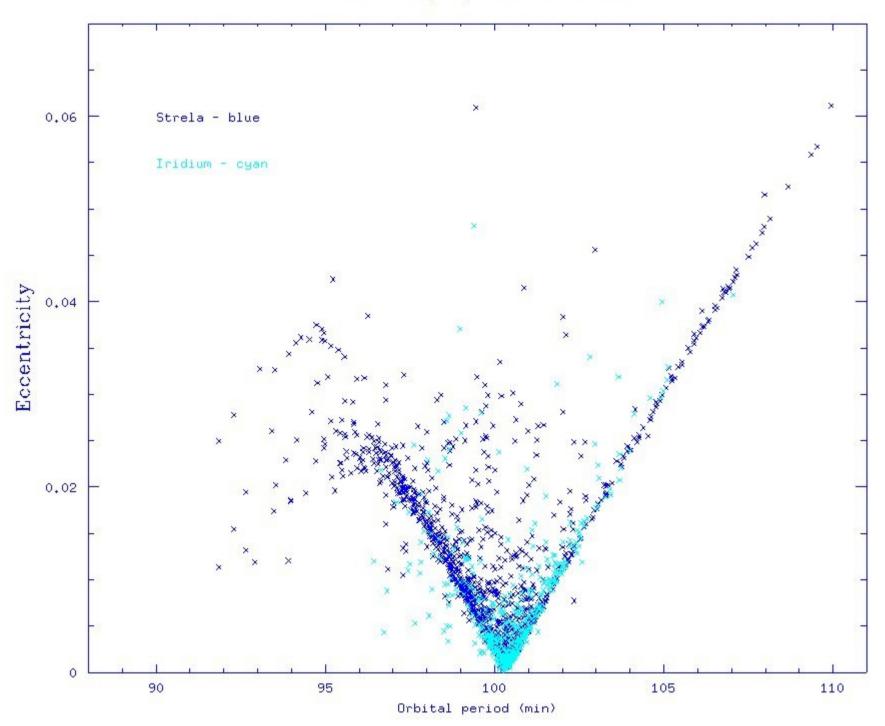
For small, light debris objects, atmospheric drag significant even at these altitudes (altitude data derived from NORAD/USSTRATCOM orbital elements via Space-Track.Org)

Current status:

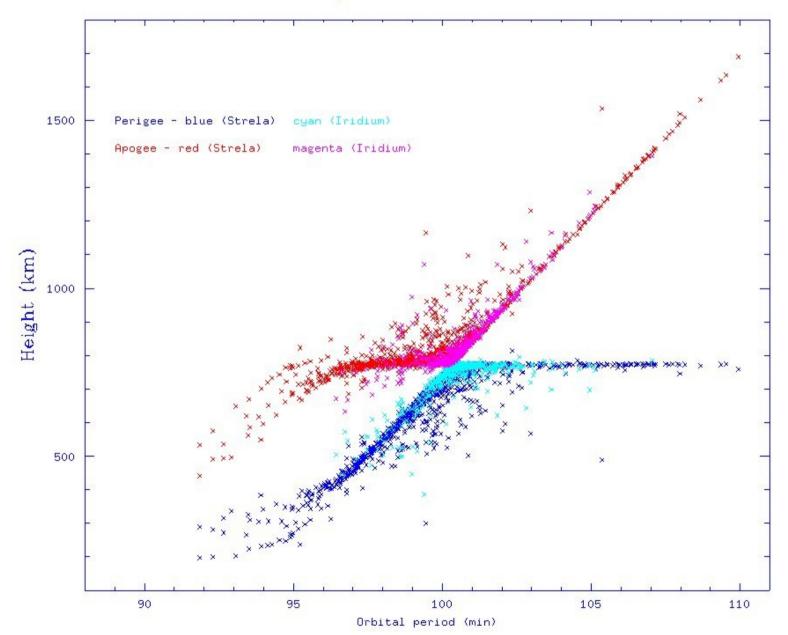
Iridium debris - 629 cataloged 286 reentered

Strela debris: 1667 cataloged 566 reentered



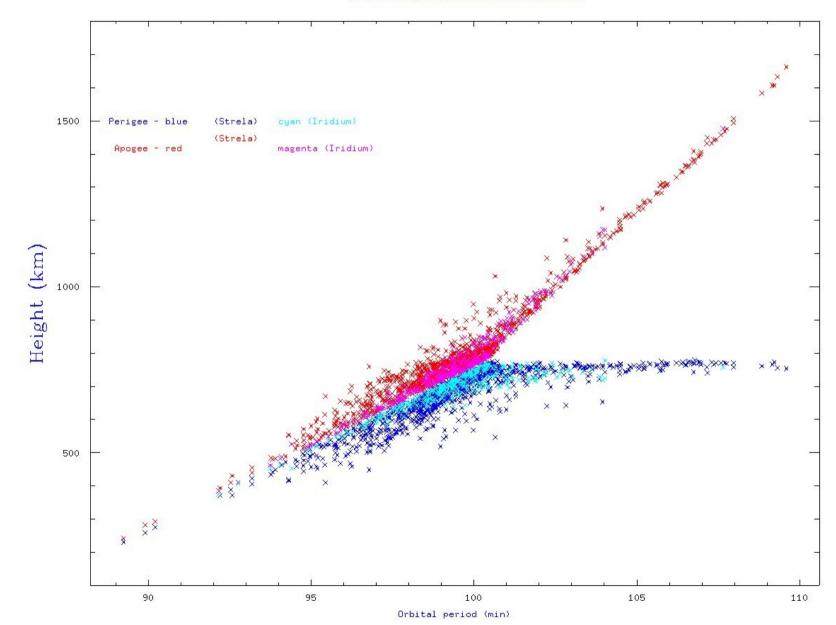


Gabbard Diagram (developed by NORAD's J. Gabbard in 1960s) shows perigee, apogee vs orbital period Can see different ejection energies of debris objects

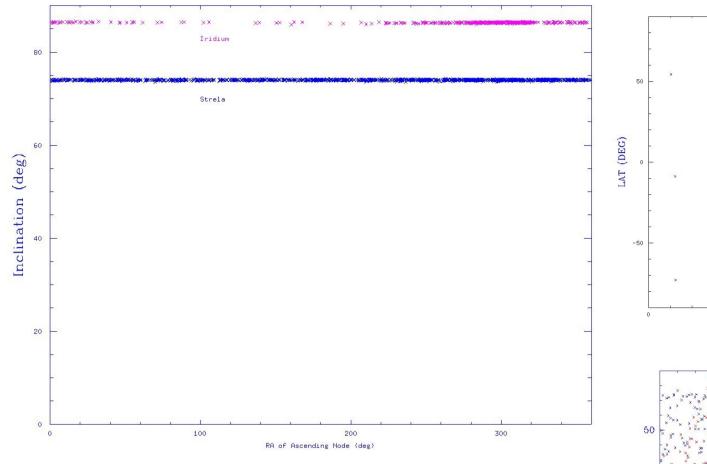


Gabbard Diagram for Iridium/Strela event - Feb 2010

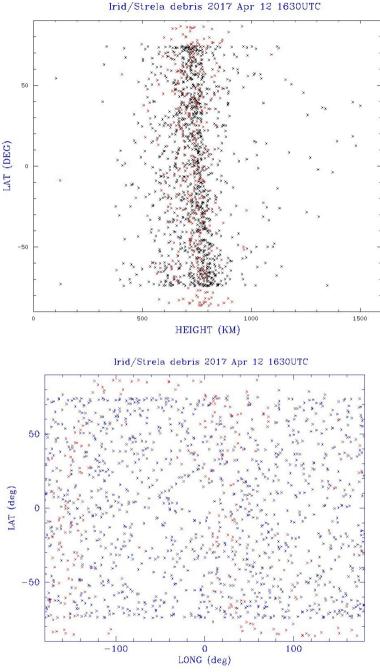
Gabbard Diagram for same event, updated to Apr 2017 Orbital decay has circularized low period objects



Gabbard Diagram from Iridium/Strela event



8 years after collision, remaining objects have spread around the Earth - all ascending node longitudes filled. But inclinations retain imprint of original two satellites.



GEO debris

GEO collisions are not so bad - r/R is 6 times larger so Keplerian velocity is sqrt(6) times smaller.

All GEO payloads going the same way round at same inclination so relative velocities are small: typically < 0.1 km/s or less for a relocating satellite

The real debris problem is from exploded rockets in GTO (Geostationary Transfer Orbit, typically 250 x 35700 km with 5 to 50 degree inclination)

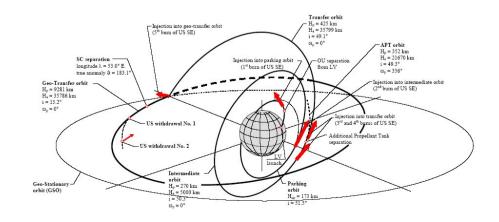
Catalog incompleteness for these is high; some data from optical telescopes, little from radars due to r**-4 dependence

Consider Falcon 9 second stage from Jan 2017 launch of Echostar satellite. Current orbit is 180 x 35767 km x 22.5 deg. Apogee velocity is 1.60 km/s

Echostar 23 payload is in 35777 x 35795 km x 0.1 deg orbit Orbital velocity is 3.07 km/s

If they collided relative velocity would be (doing the vector math) 1.7 km/s

- 47 times less KE/kg than the Iridium-Strela event

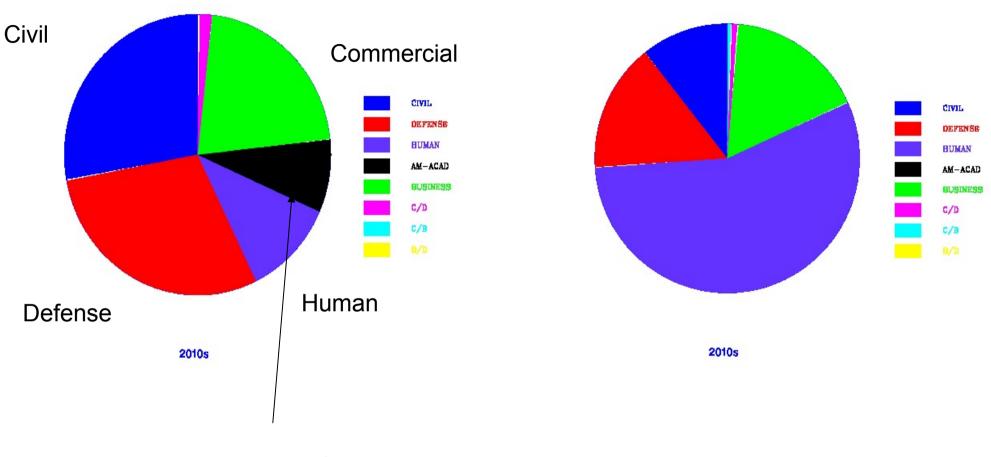


Scheme of SC Injection into Target Orbit from 1st Ascending Node of Parking Orbit

Proton launch profile for Turksat-4B (Khrunichev)

Satellite Classes

Satellite Tonnage



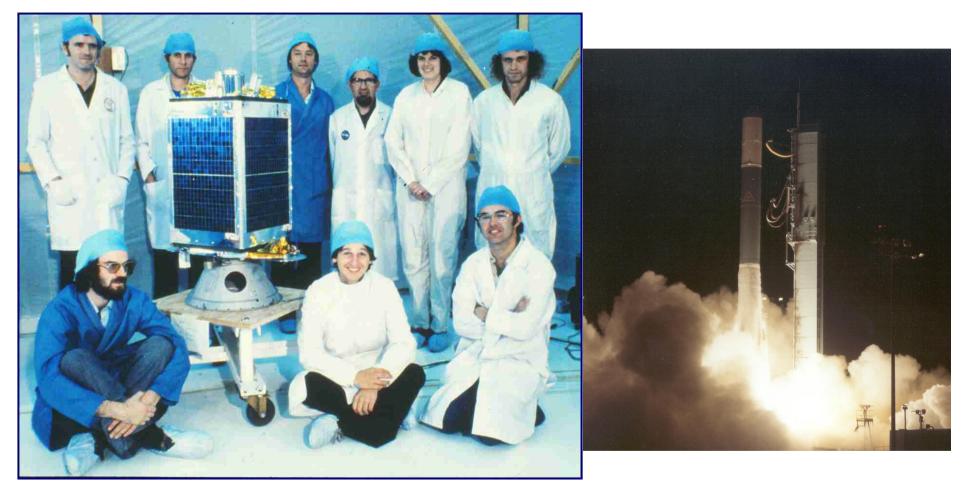
Non-profit

Lump all countries together – division between military, civilian and commercial is about even if you exclude human spaceflight (most of the tonnage, and money)

Non-profit an important sector by number of satellites, but tonnage is negligible

Guildford, 1981: University of Surrey team (under Martin Sweeting) builds amateur radio satellite UoSat-1

It becomes the basis of a series of cheap commercial satellites affordable by developing countries





Alsat (Algeria) 2002

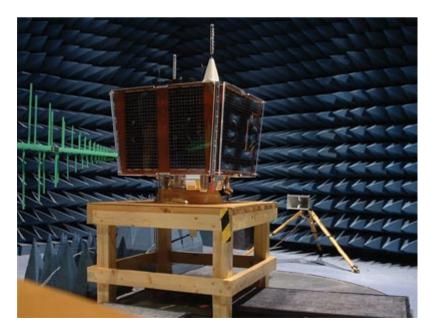


Tiungsat (Malaysia) 2000



Fasat (Chile) 1998

Posat (Portugal) 1993





Uribyol S Korea 1992

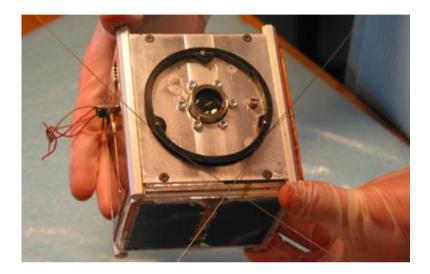


PoSAT-1

Nigeriasat-2 2011

Bilsat (Turkey) 2003

Cubesats: 1 kg, 10 cm (2 lb, 4 in for the metric impaired) Standard kit for universities to make students build sats in engineering courses Can also make '3U' cuboids 30 x 10 cm 97 Cubesats launched 2003-Feb 2013 by 66 organizations in 20 countries



Aalborg U. 2003



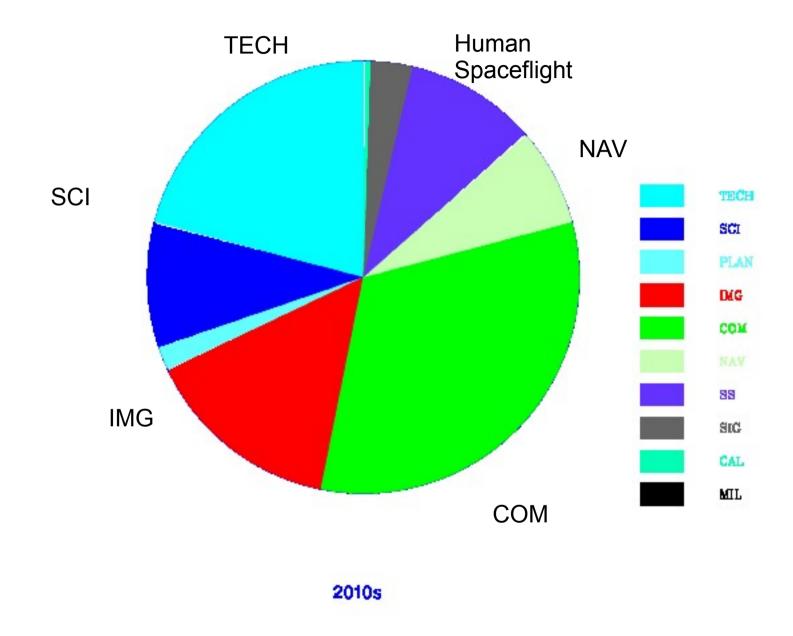
Cubesat deploy from ISS, 2012

Univ. of Tokyo, 2003



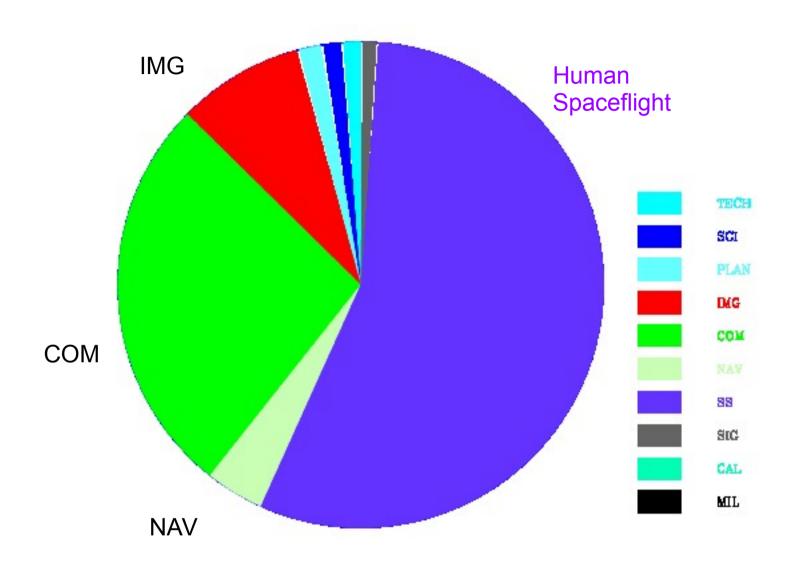
Triple-cube Quakesat, Stanford 2003

Satellite Categories



In the 2010s, most sats are either communications or imaging; technology development (including student satellites) also a big sector

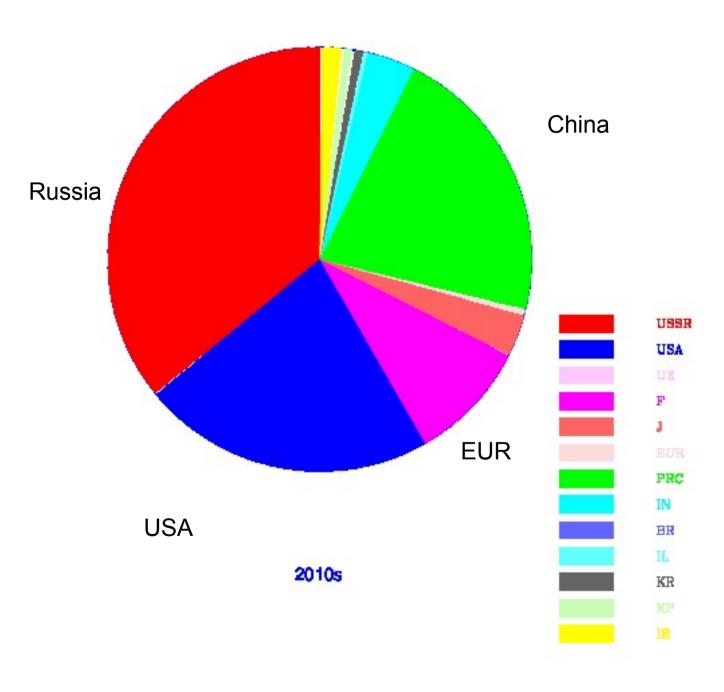
Satelite Tonnage (including human spaceflight)



2010s

By mass however, human spaceflight dominates – comms still next Tech/student satellites vanish, they are mostly little cubesats which don't weigh much 4-yr total 900 t robotic, 1100 t 6 x Shuttle + ISS/PRC

Globalized Space Launch Capability



Today the space launch market has many more players

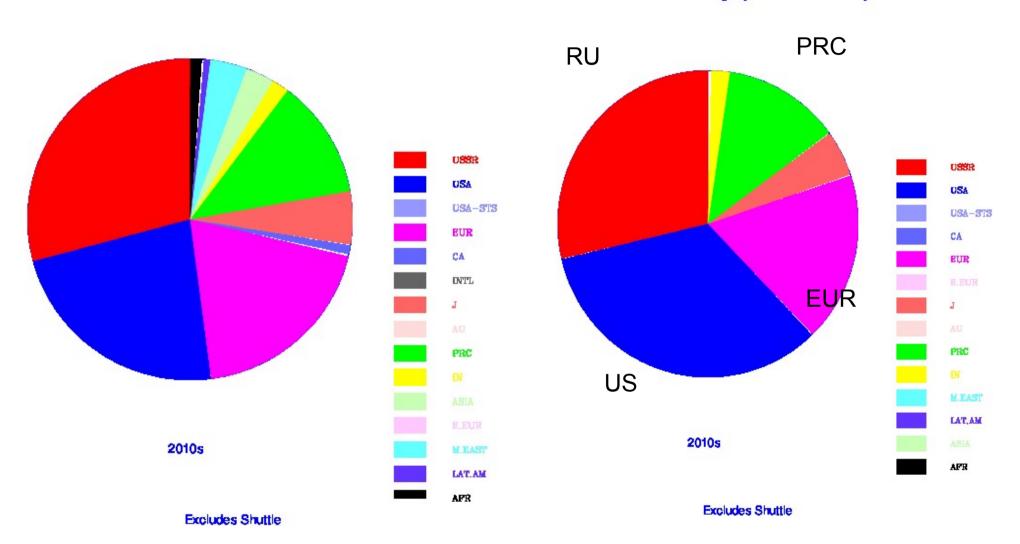
In 2012 China had as many orbital launch attempts as the US

12 countries plus ESA/Arianespace have launched satellites; Brazil has also tried but failed.

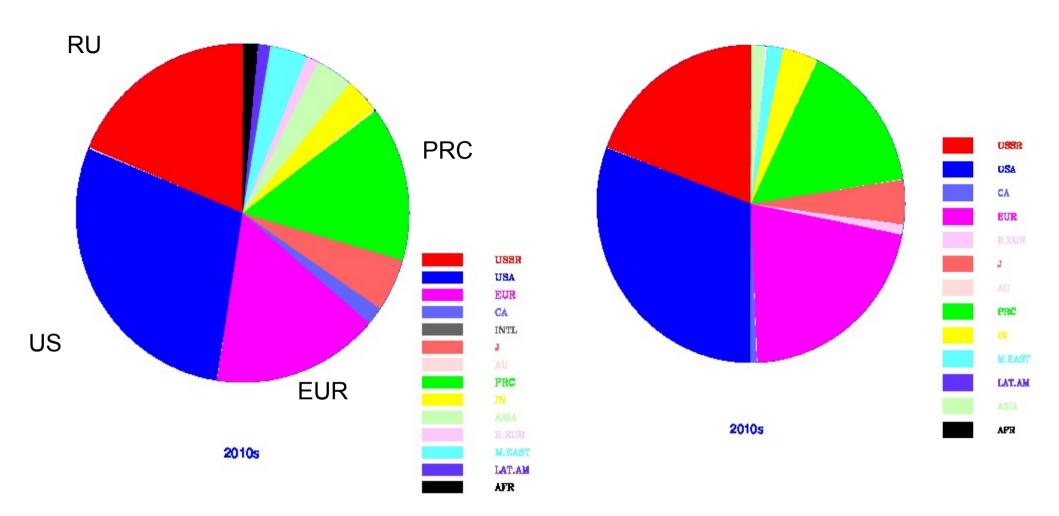
North and South Korea are the latest members of the club

Satellite Tonnage by Owner Country

Satellite Tonnage by Manufacturer Country



The 'other' countries almost vanish when considering tonnage - Their satellites are usually tiny cubesats



Lots of countries OWN satellites – too many to show on the chart, so I grouped together E. Europe (pink), Africa (black), Latin America (dark purple), and Asiaother-than-China/India/Japan (light green)

Russia, US, W. Europe and China dominate; next Japan (orange), and India (yellow)